

Data Assimilation Changes Low-Level Jets in Mars GCMs

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(How much) should we believe GCMs?

- They work ... as far as we can see
- Limited non-temperature dynamics obs
 - Can reanalysis be a substitute?
- **Today: How good are low level flows?**
 - Diagnostic work, then theory

$\tau = 0.94$

2.9

4.1

3.8

4.7

Motivation, datasets, and method

- Basic dynamics, dust cycle, lander ops

1205
11:14

1220
11:04

1225
11:30

1233
10:55

1235
10:53

Opportunity Sol Number and Local True Solar Time

$\tau = 0.94$

2.9

4.1

3.8

4.7

Motivation, datasets, and method

- Basic dynamics, dust cycle, lander ops
- Two reanalyses and their free runs
 - MACDA—UK GCM, **analysis correction**
 - EMARS—GFDL GCM, **EnKF**
- Both use TES obs

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Opportunity Sol Number and Local True Solar Time

$\tau = 0.94$

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Motivation, datasets, and method

- Basic dynamics, dust cycle, lander ops
- There's ultimately only one truth so ...
 - Two “good” reanalyses should agree
 - GCMs should agree w/ “good” reanalyses

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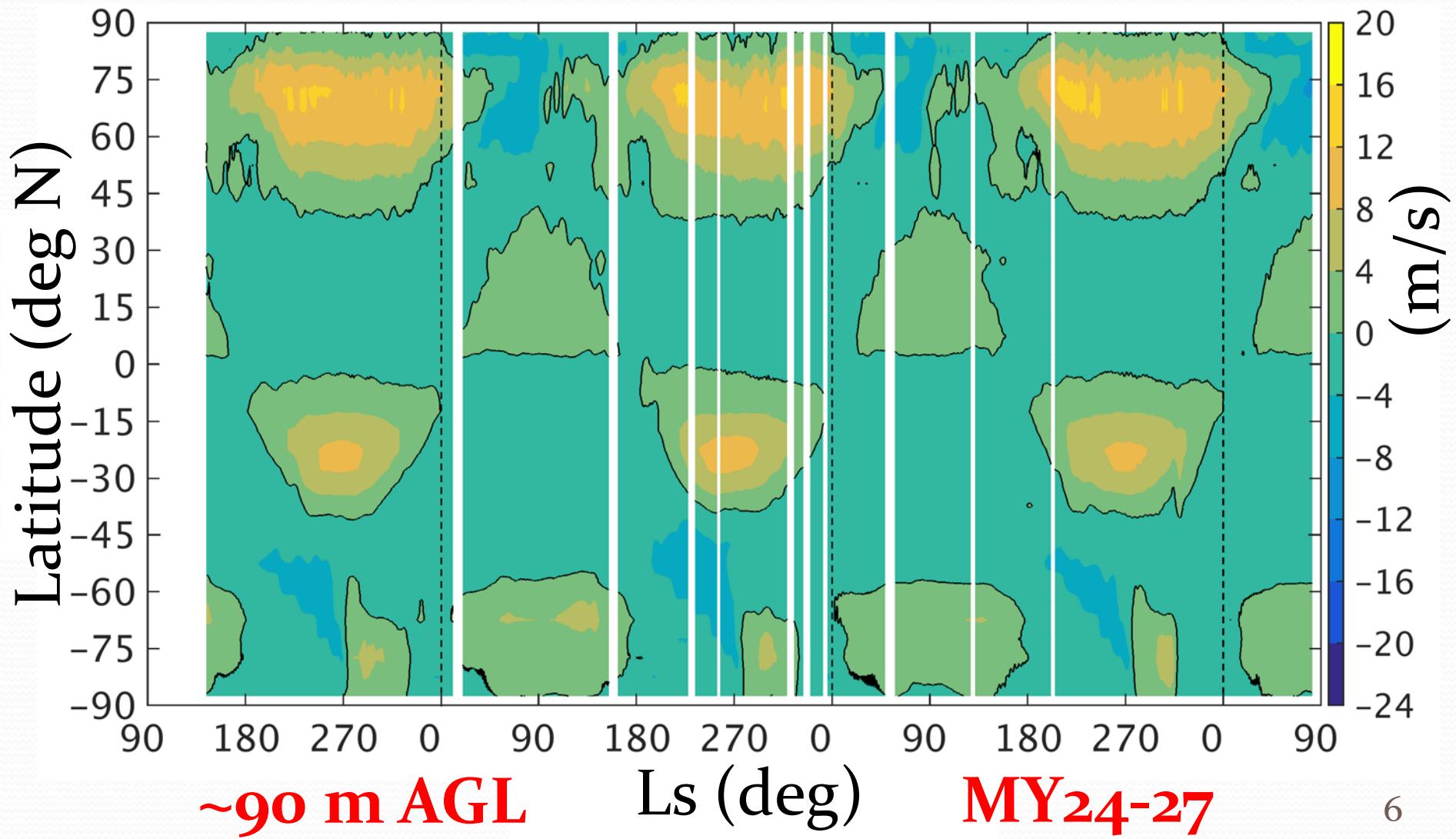
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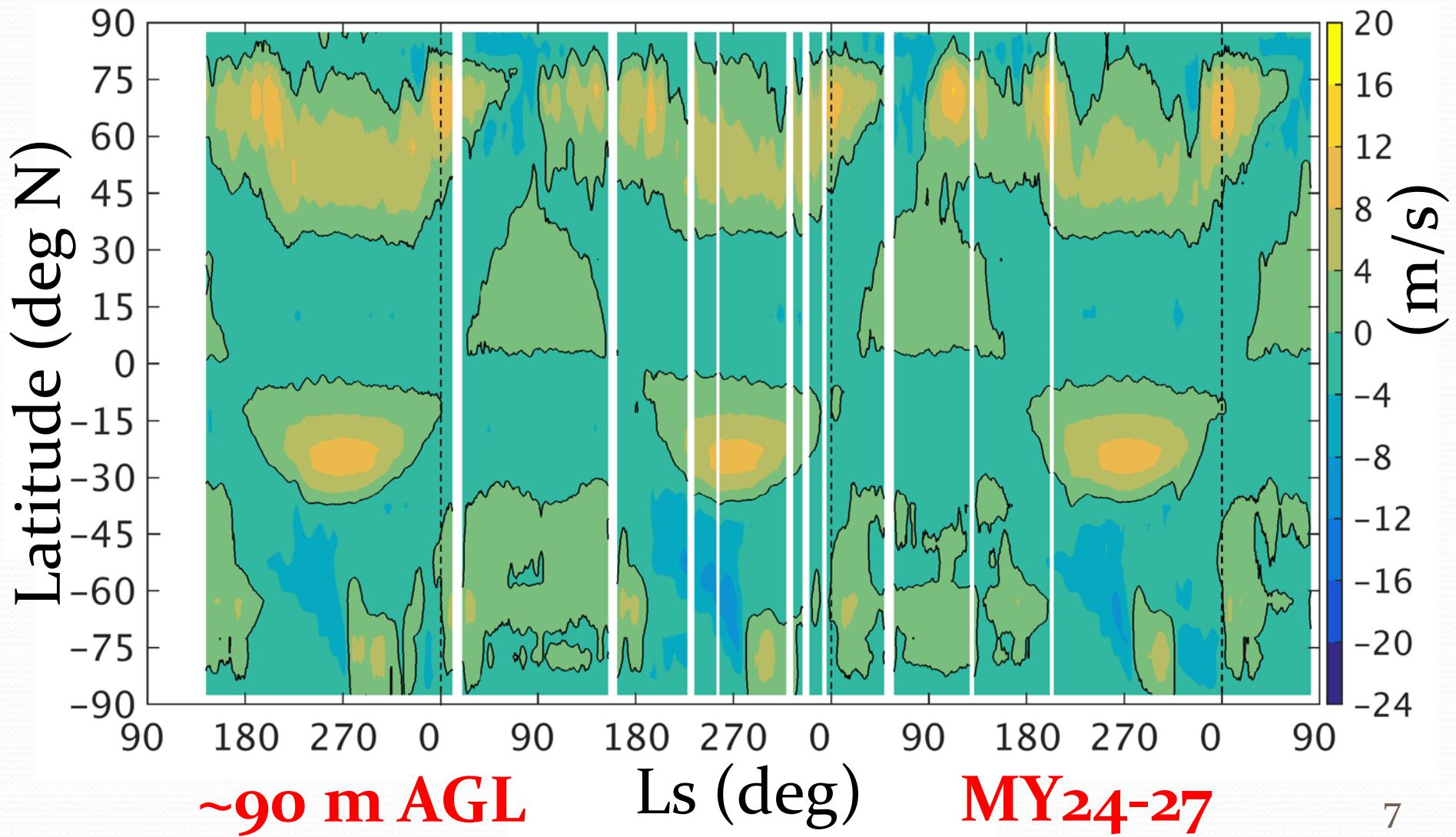
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Opportunity Sol Number and Local True Solar Time

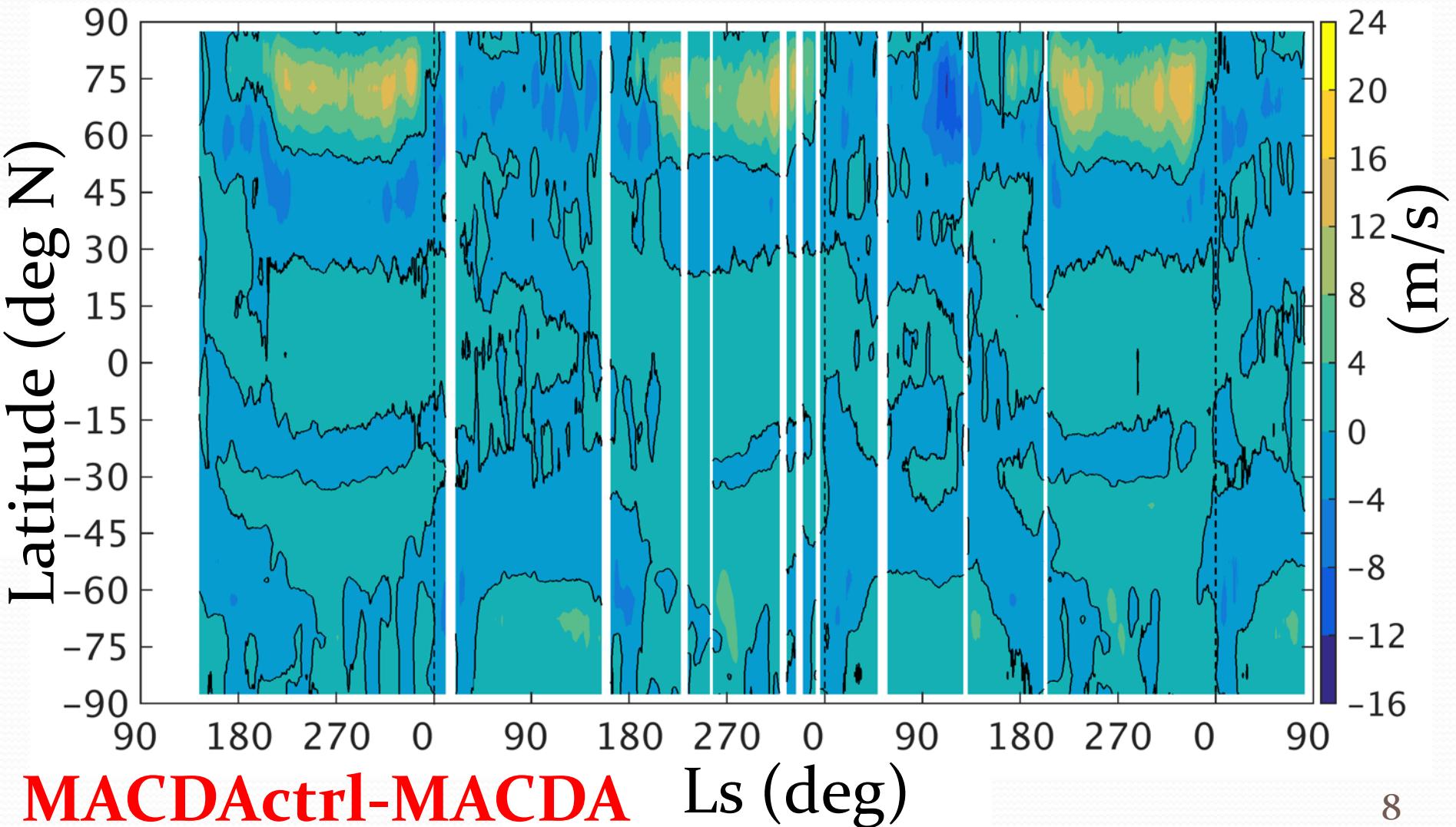
Strong NH winter jets (MACDActrl)



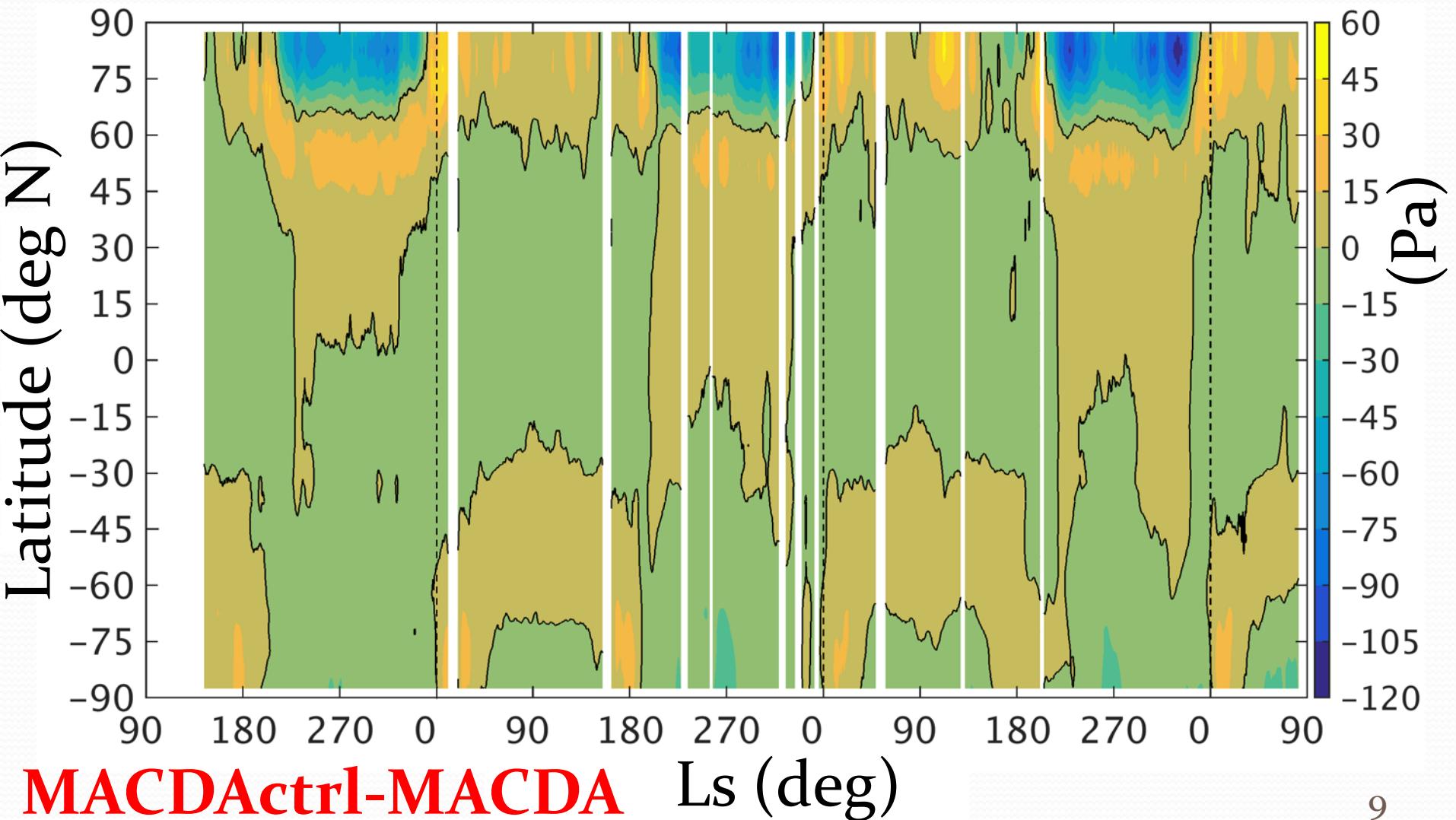
MACDA jets weaker, farther south



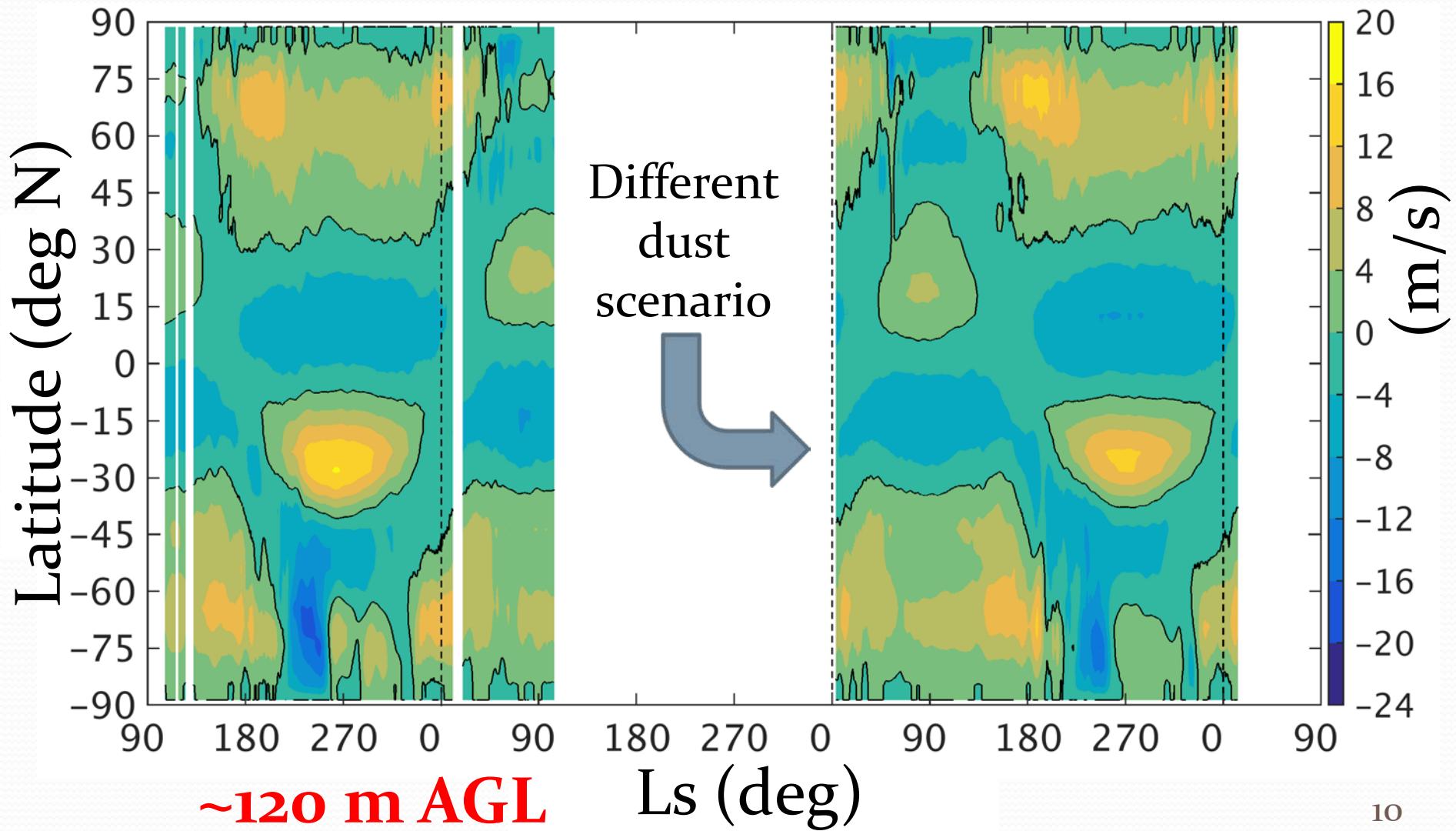
DA effect is interannually repeatable



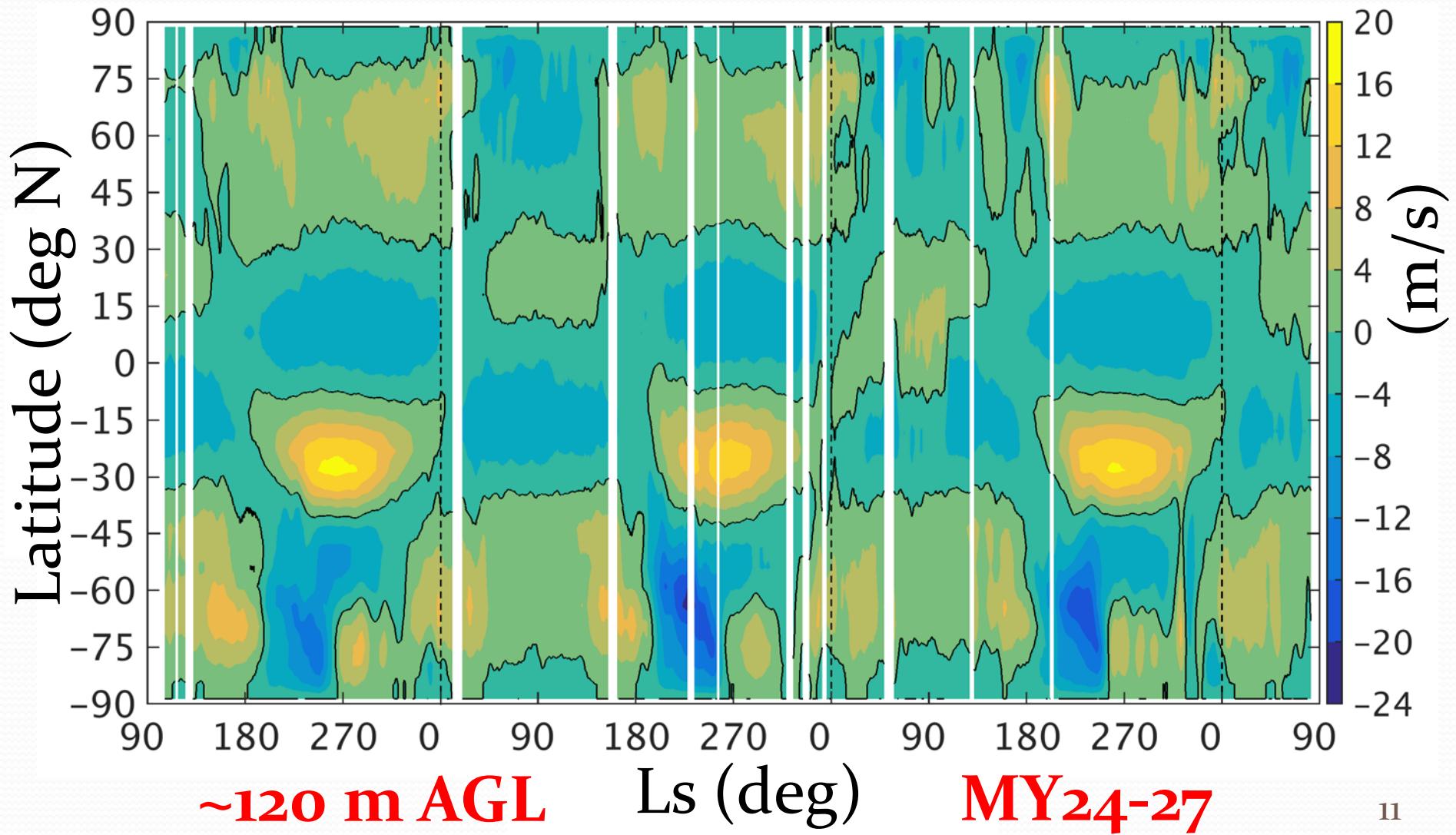
DA (relatively) raises polar p_{surf}



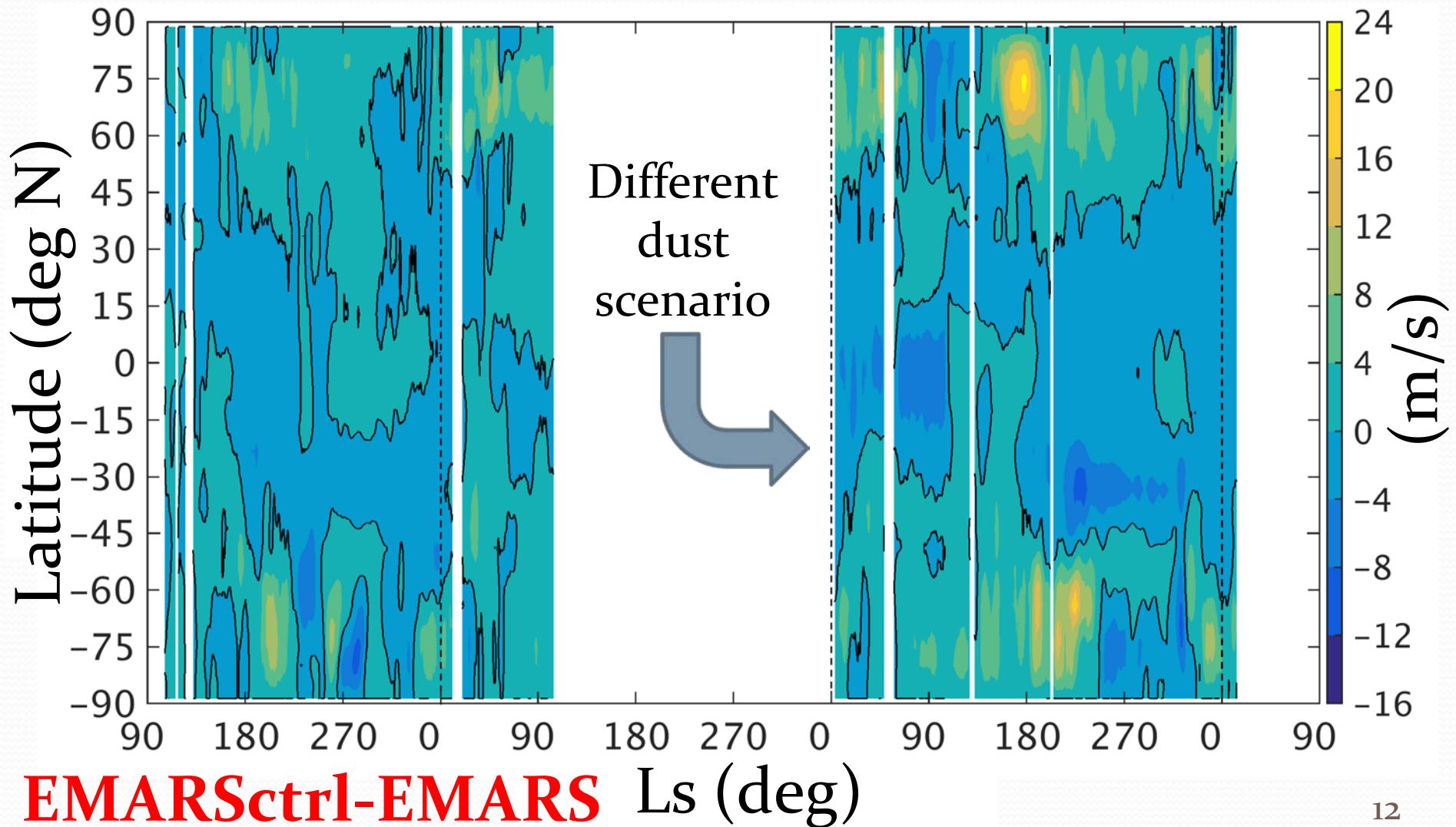
Similar seasonal cycle in EMARSctrl



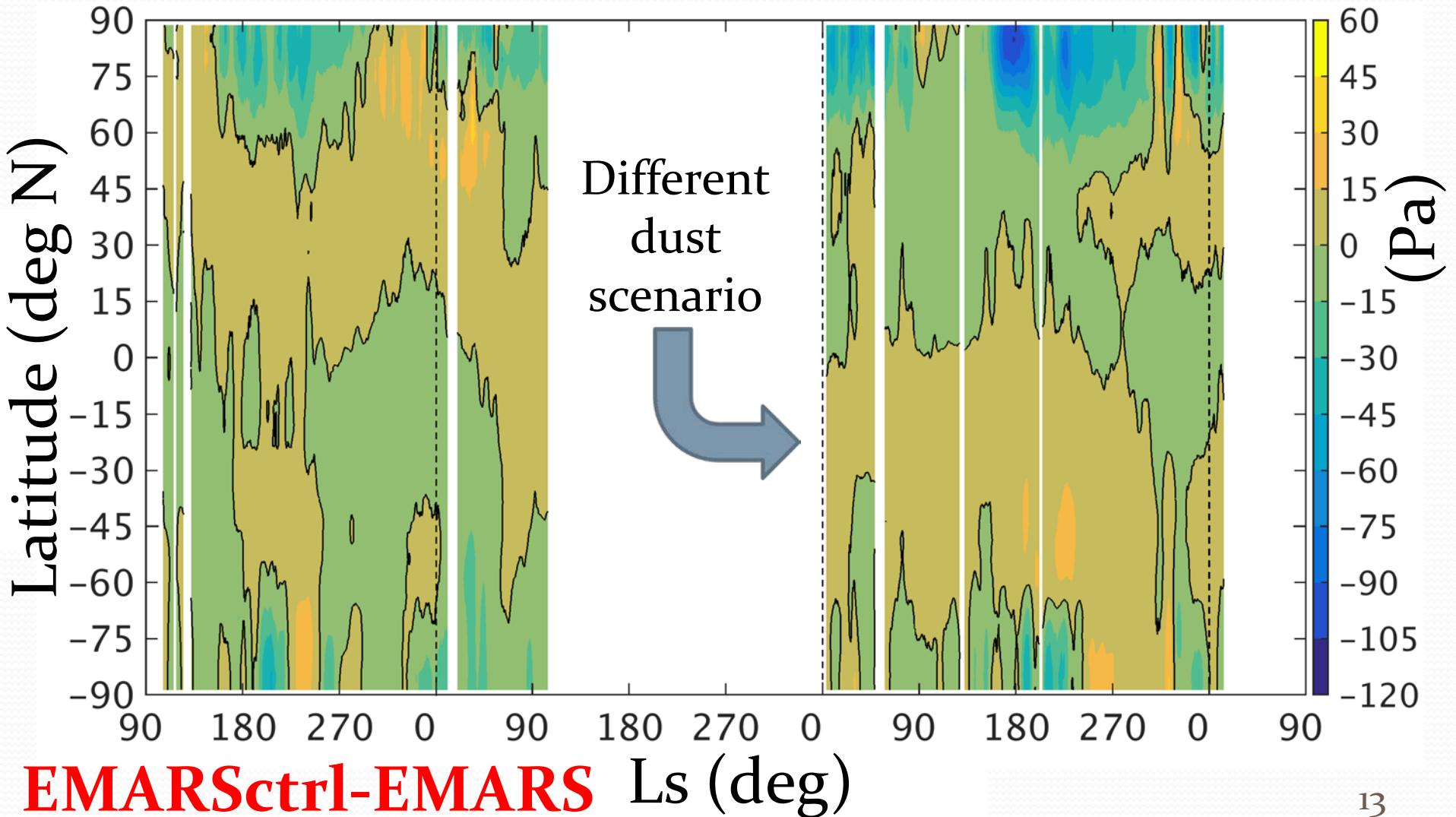
DA weakens the EMARS jet



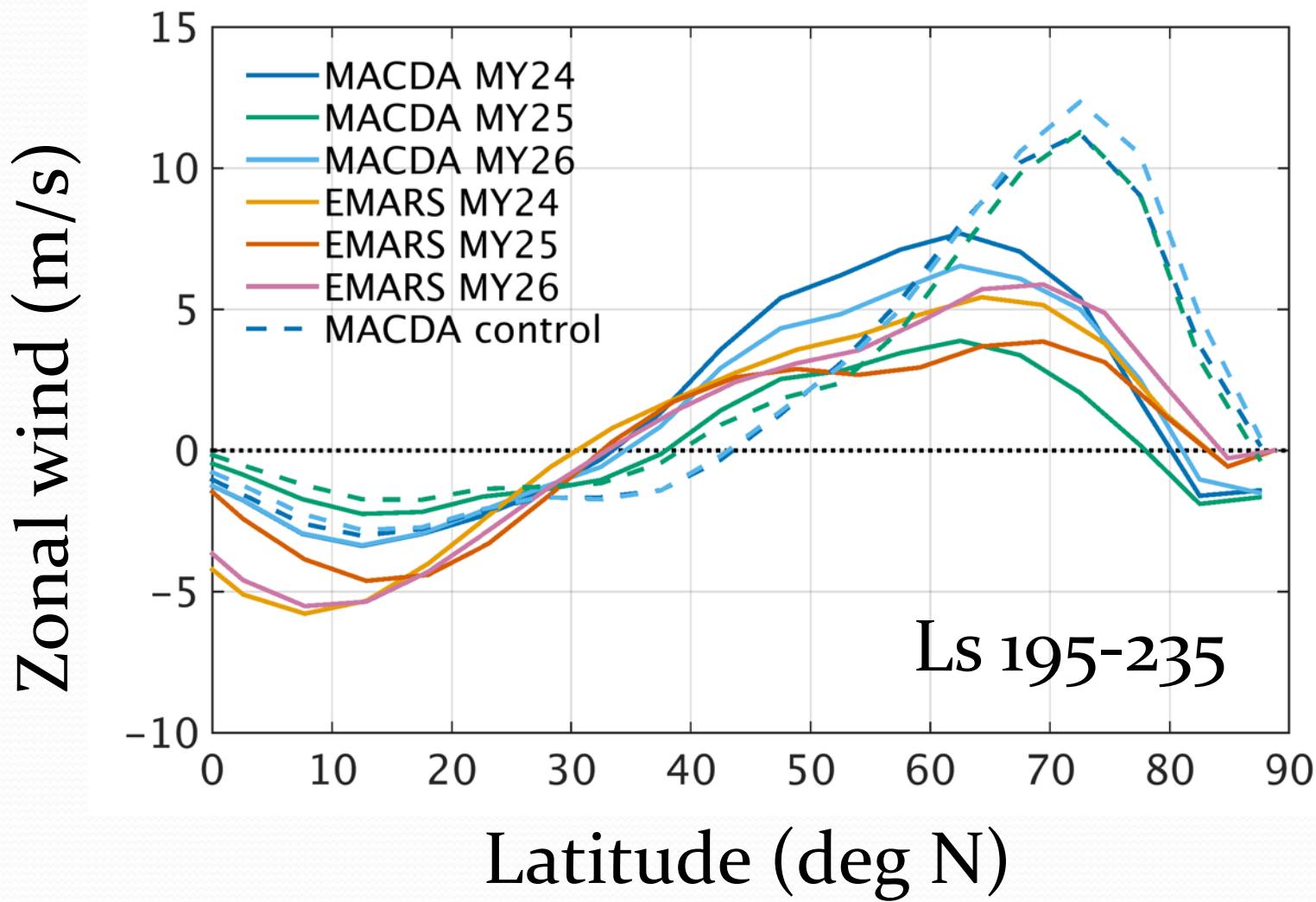
DA response less than in MACDA



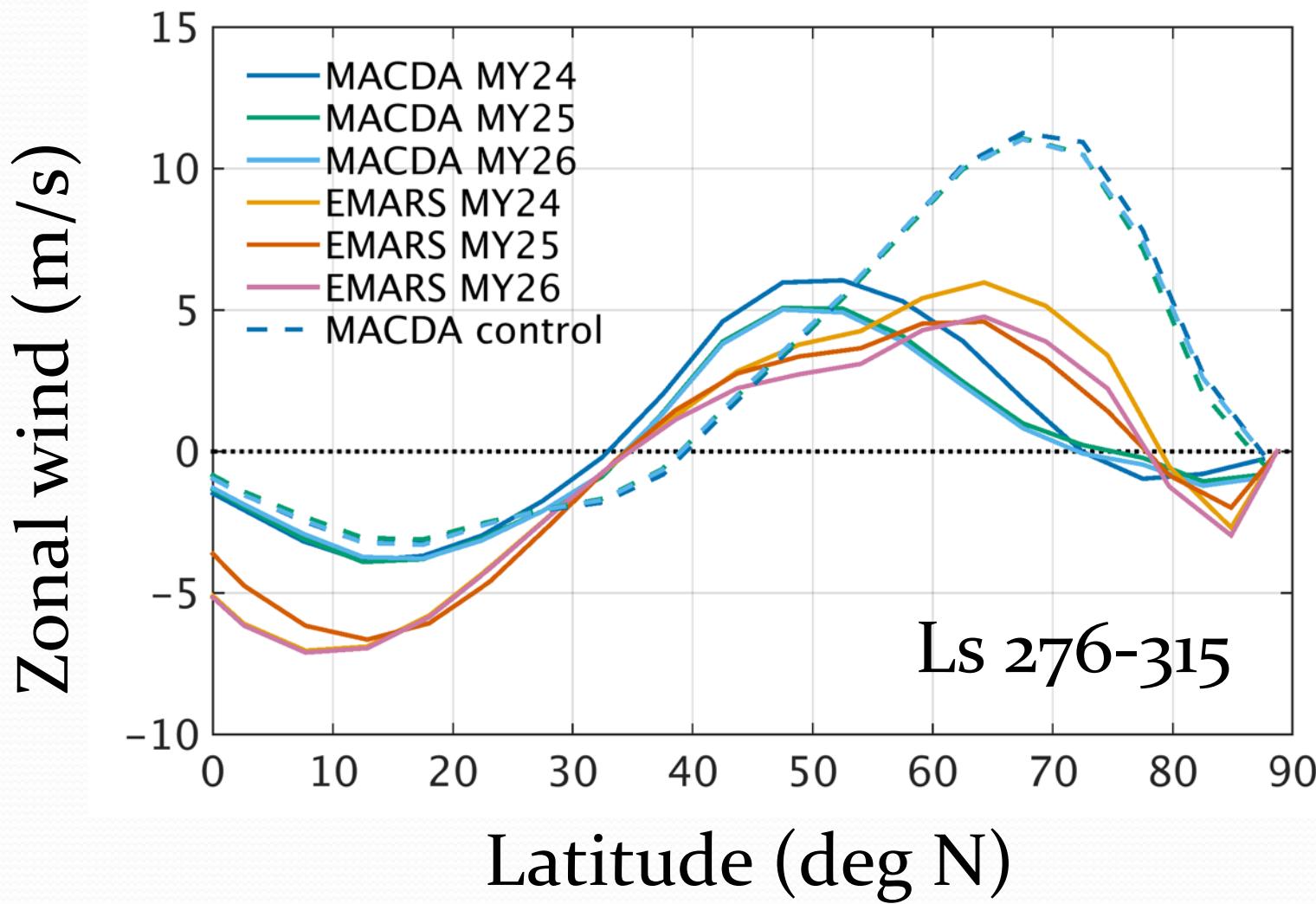
DA effect on p_{surf} phased differently



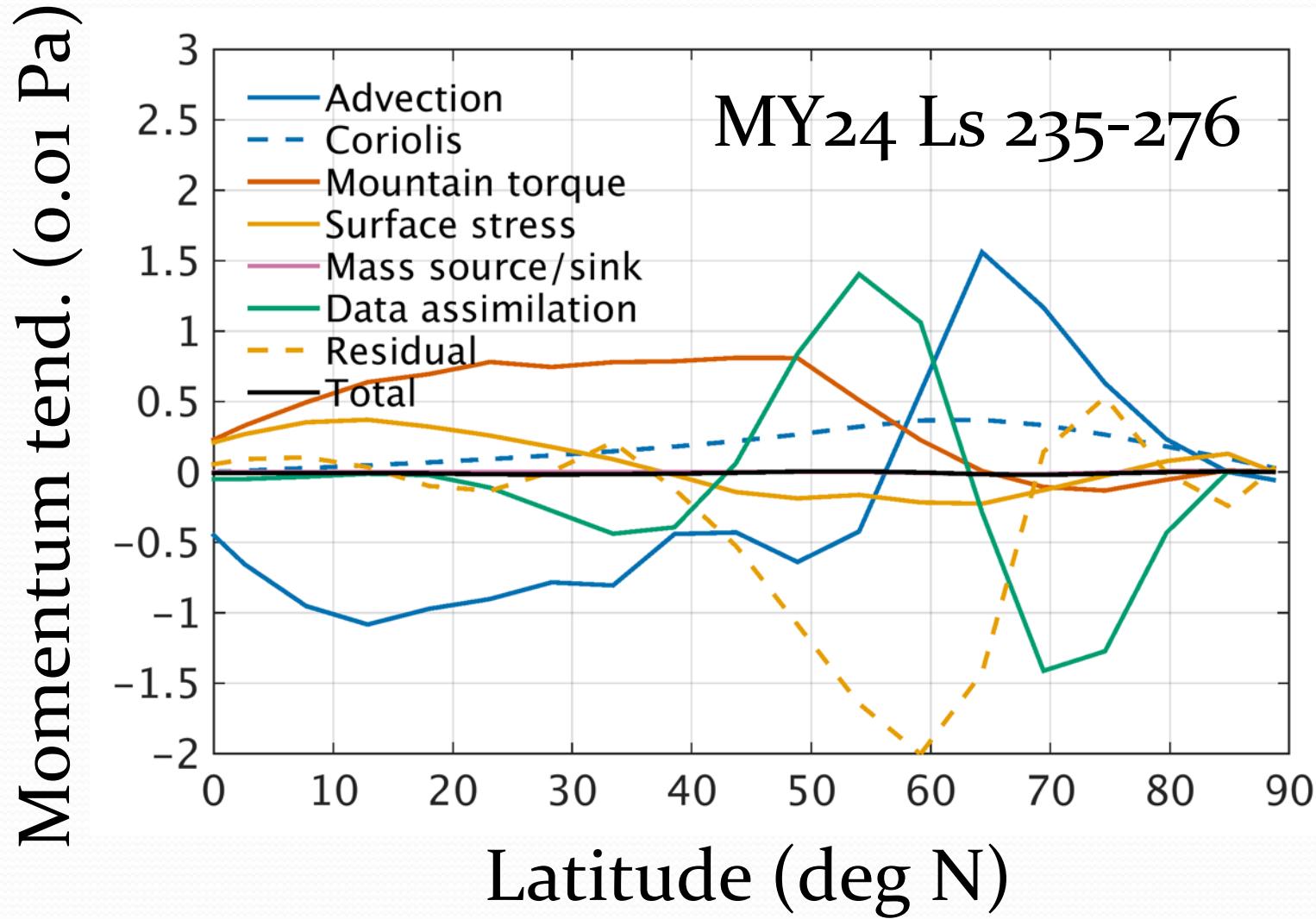
DA adds variability to MACDA



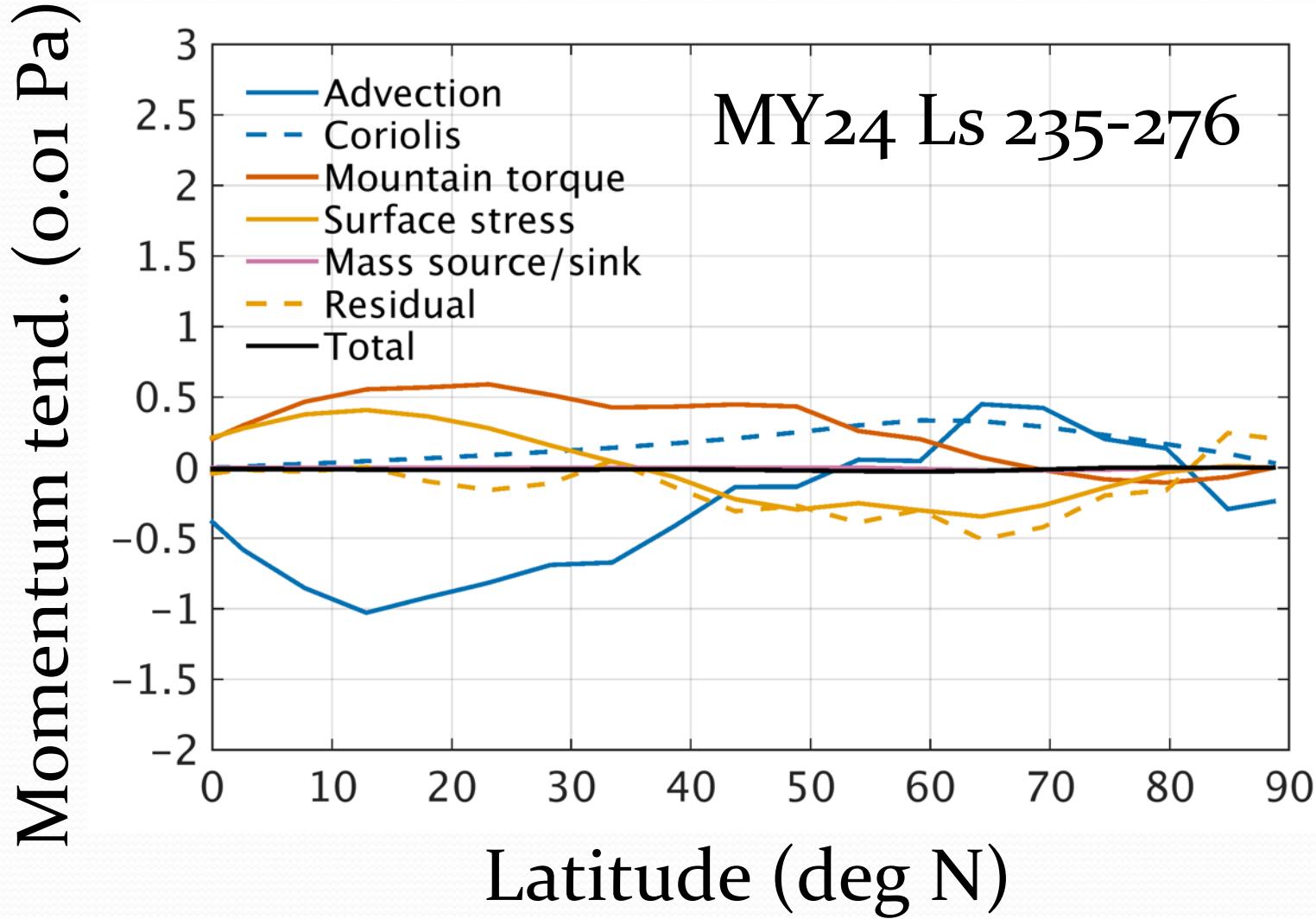
Jet positions can disagree



DA tendencies directly shift jet



Warning: budget closure issue



Summary of diagnostic work

- T assimilation
 - Weakens NH jet
 - Shifts it equatorward, esp. in MACDA
 - Effect smaller in EMARS than MACDA
- Surface jet position—a **common GCM bias?**
- **But** reanalyses still clearly disagree

Theoretical aside

- In a toy problem (Boussinesq QG theory), perfect 4D temperature data does **NOT** uniquely determine flow
- Also need heating/friction info ... physics

Broader conclusions/speculations

- What other GCM errors can't we see?
- Need good models for data assimilation
 - Budget closure as evaluation metric?
- Better near-surface obs might help
 - Aeolus or RO constellation?

Some theory

- What can we learn from temperature alone?
- Toy problem: Boussinesq QG theory
- **Will show:**
 - Must know heating & friction to deduce surface flow

Some theory

- What can we learn from temperature alone?
- Toy problem: Boussinesq QG theory

Momentum

$$\frac{D\vec{u}}{Dt} + f\hat{z} \times \vec{u} = -\frac{\nabla p}{\rho_{00}} + \vec{F}$$

Mass conservation

$$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \frac{\partial w}{\partial z} = 0$$

Thermo

$$\frac{D\rho}{Dt} = -S$$

Hydrostatic balance

$$\frac{\partial p}{\partial z} = -g\rho$$

Boussinesq QG equations

QGPV equation

$$\frac{D_g q}{Dt} = \nabla \times \vec{F} + \frac{gf_0}{\rho_{00}} \frac{\partial}{\partial z} \left[\frac{S}{N^2} \right]$$

QGPV definition

$$q = \beta y + \left[\nabla^2 + \frac{\partial}{\partial z} \frac{f_0^2}{N^2} \frac{\partial}{\partial z} \right] \varphi$$

Winds $(u, v) = \left(-\frac{\partial \varphi}{\partial y}, \frac{\partial \varphi}{\partial x} \right)$

“Temperature” $\frac{\partial \varphi}{\partial z}$

Boundary conditions @ $z = 0$ & H

$$\frac{D}{Dt} \frac{\partial \varphi}{\partial z} = \frac{g S}{f_0 \rho_{00}}$$

The big idea

Assume solution

$$\varphi(x, y, z, t), \vec{F}, S$$

Then $\varphi(x, y, z, t) + \Delta\varphi(x, y, t), \vec{F} + \Delta\vec{F}, S + \Delta S$

has different sfc flow but same “temperatures”

Subtract solutions, w/ $\Delta\varphi \neq 0$ can find $\Delta\vec{F}$ & ΔS

E.g.
$$\left[-\frac{\partial \Delta\varphi}{\partial y} \frac{\partial}{\partial x} + \frac{\partial \Delta\varphi}{\partial x} \frac{\partial}{\partial y} \right] \frac{\partial \varphi}{\partial z} = \frac{g \Delta S}{f_0 \rho_{00}}$$

at boundaries

The big idea

Assume solution

$$\varphi(x, y, z, t), \vec{F}, S$$

Then $\varphi(x, y, z, t) + \Delta\varphi(x, y, t), \vec{F} + \Delta\vec{F}, S + \Delta S$

has different sfc flow but same “temperatures”

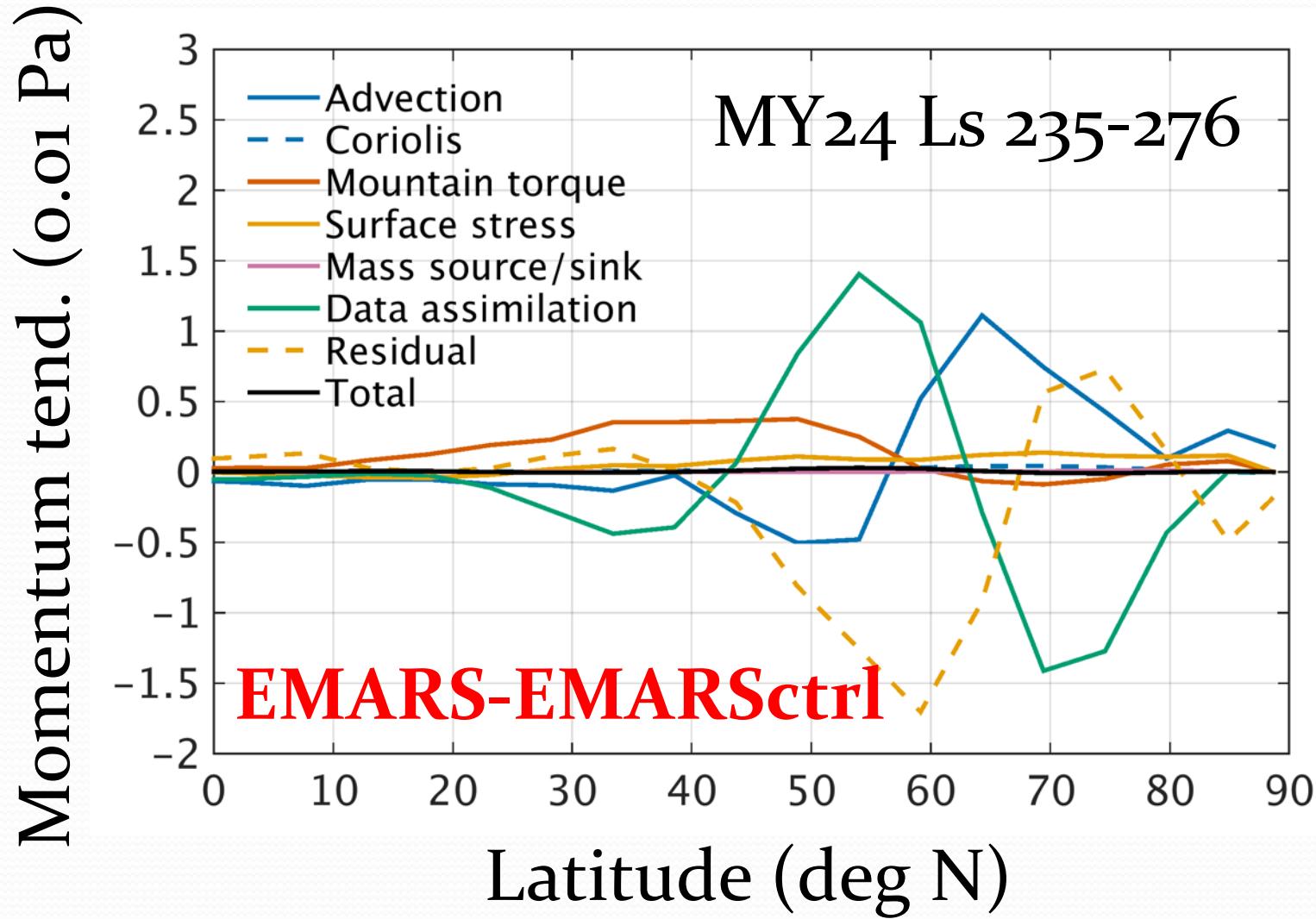
Subtract solutions, w/ $\Delta\varphi \neq 0$ can find $\Delta\vec{F}$ & ΔS

Must know \vec{F} and S to get barotropic φ

Caveats and implications

- Heating/friction not totally unconstrained
- Perturbed physics for ensemble DA = good?
- Problem for parameter estimation?

Momentum budget difference



Are GCM & reanalysis results robust?

- Aside: should DA work near surface?

	MACDA	EMARS
T data	TES	TES
Dust data	TES	TES
Control	Yes	Yes
GCM	UK spectral	GFDL lat-lon FV
DA method	Analysis correction	Ensemble Kalman filter

Mars low-level circulation

- Almost all Mars atmospheric dynamics data is infrared temps from orbiters
- What about wind, etc.?
 - Use temps to constrain Mars GCM
 - Model + observations = reanalysis
- How well does this work at low altitudes?

The big idea

Assume solution

$$\varphi(x, y, z, t), \vec{F}, S$$

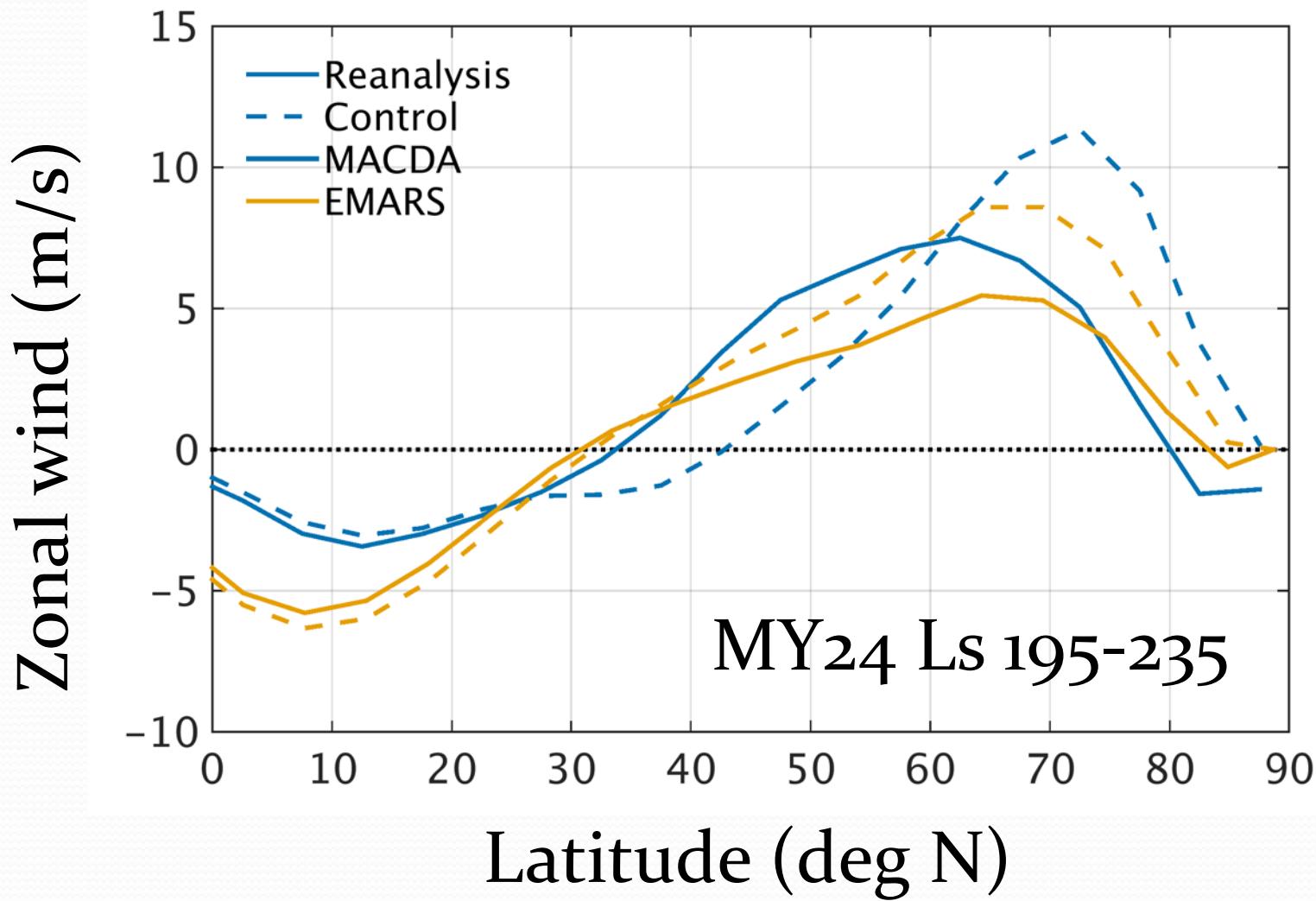
Then $\varphi(x, y, z, t) + \Delta\varphi(x, y, t), \vec{F} + \Delta\vec{F}, S + \Delta S$

has different sfc flow but same “temperatures”

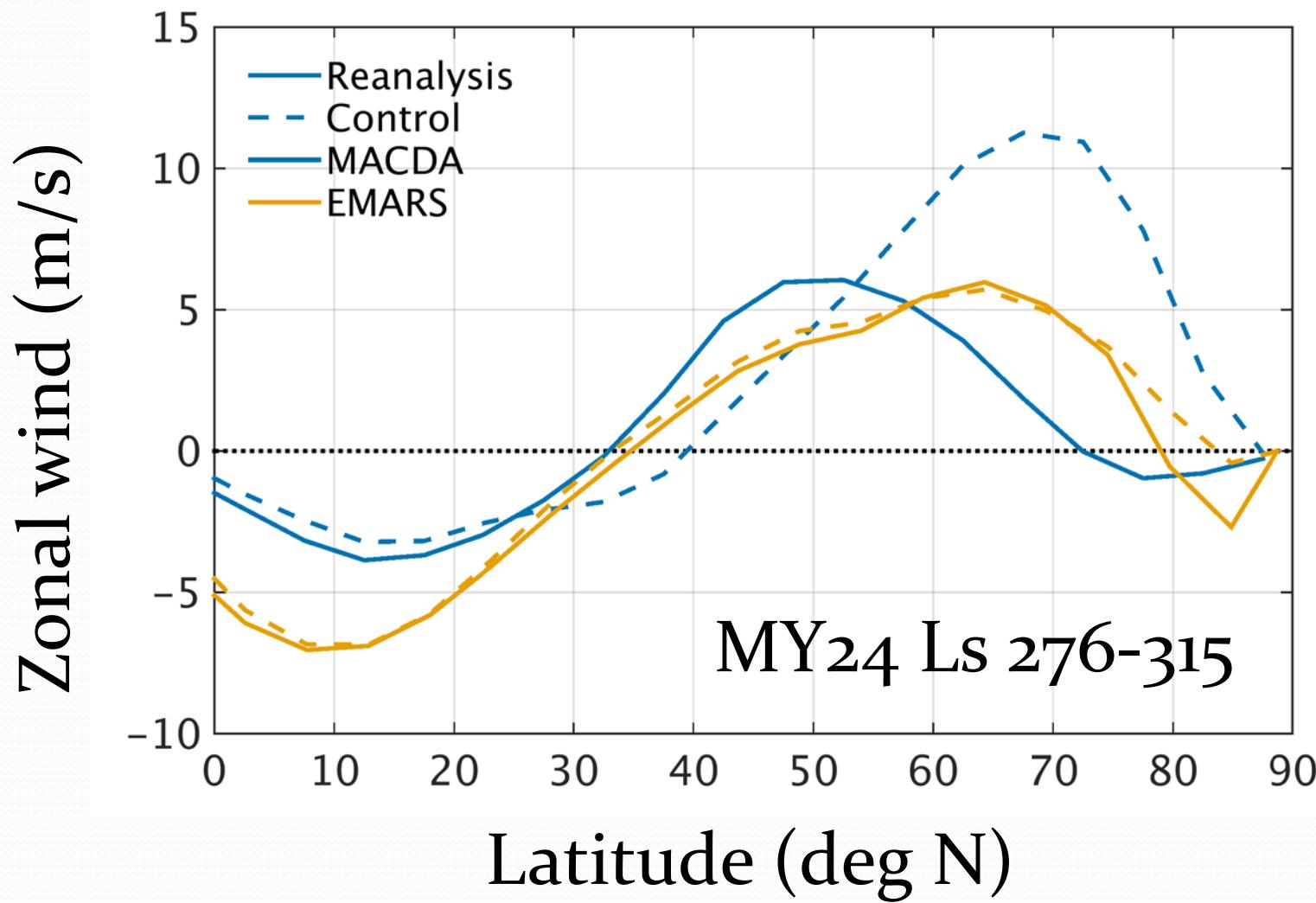
Given $\Delta\varphi \neq 0$ can find $\Delta\vec{F}$ & ΔS

Must know \vec{F} and S to get barotropic φ

No shift in EMARS



Almost no DA effect in EMARS



\sim 1 km zonal wind in MACDA reanalysis

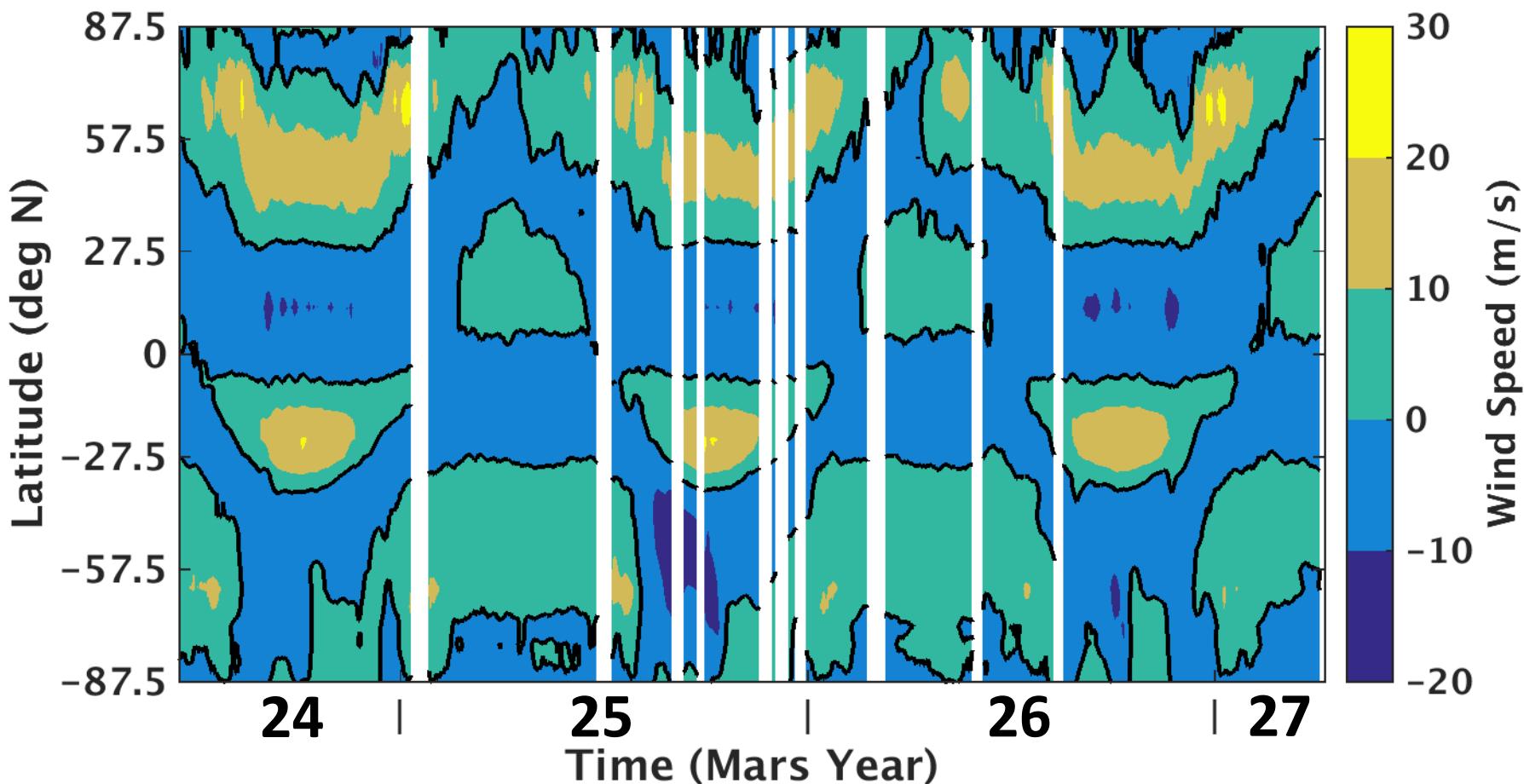


Figure: G. Davis

\sim 1 km zonal wind in control run

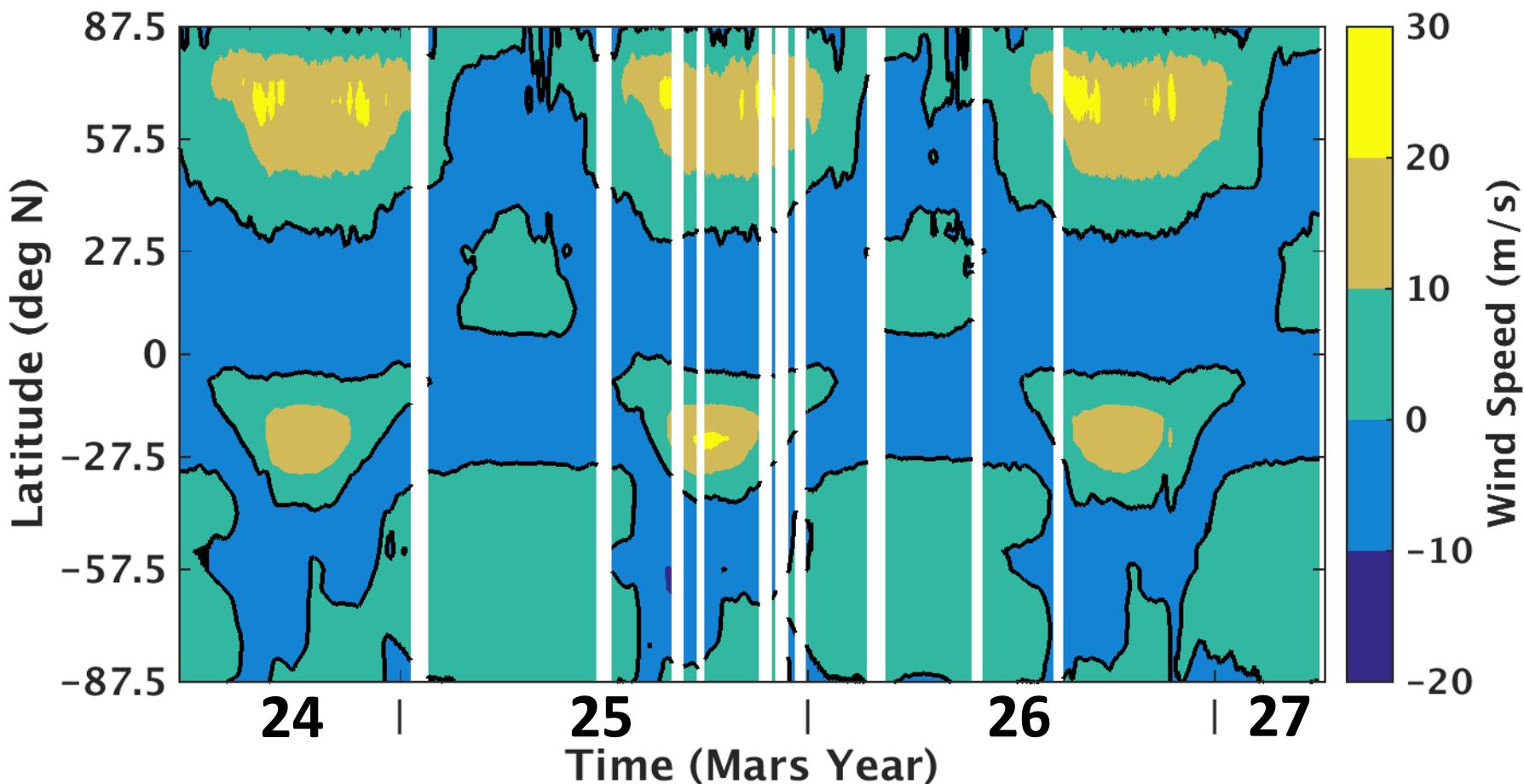


Figure: G. Davis

Difference (control – reanalysis)

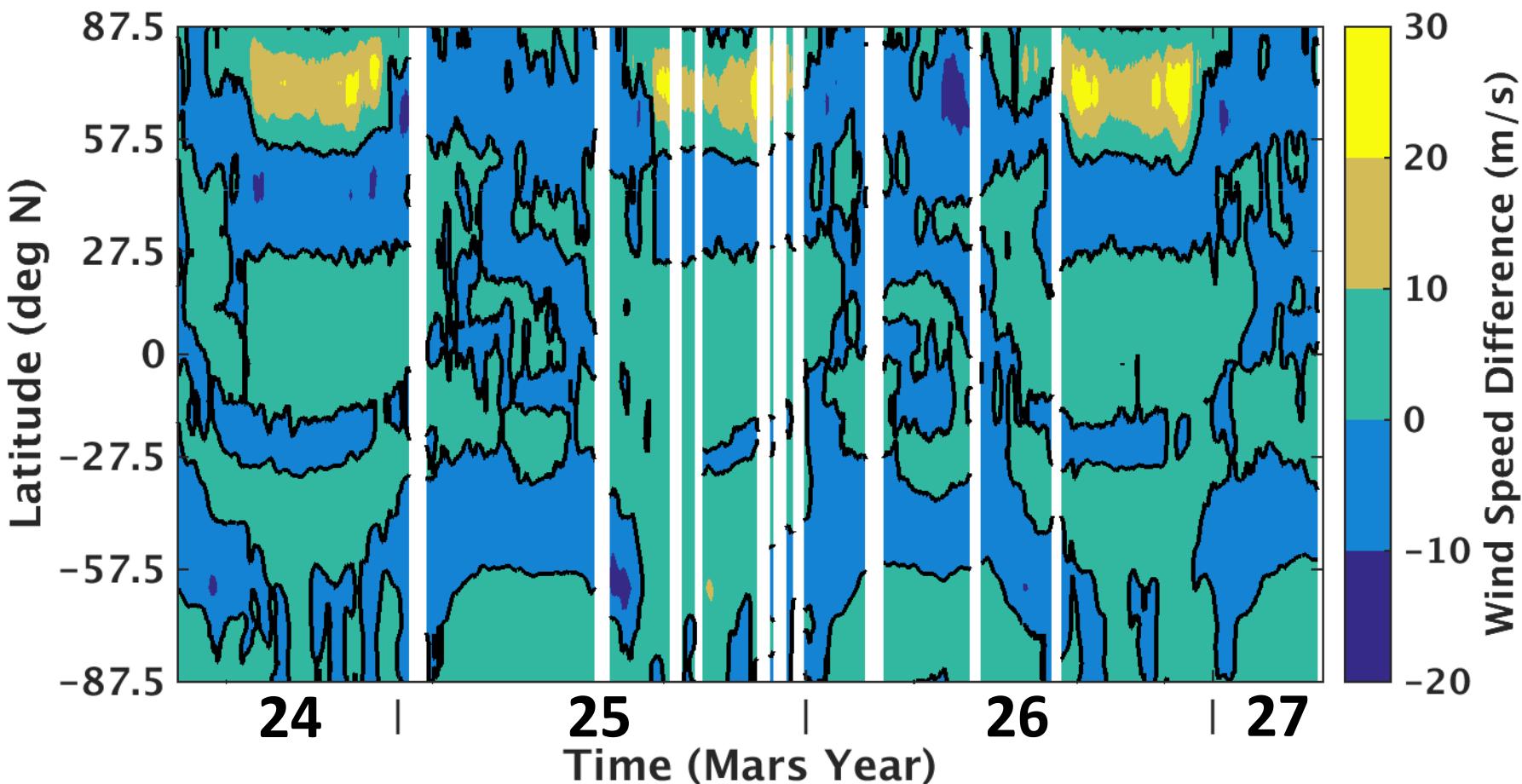


Figure: G. Davis

P_{surf} & zonal jet differences are linked

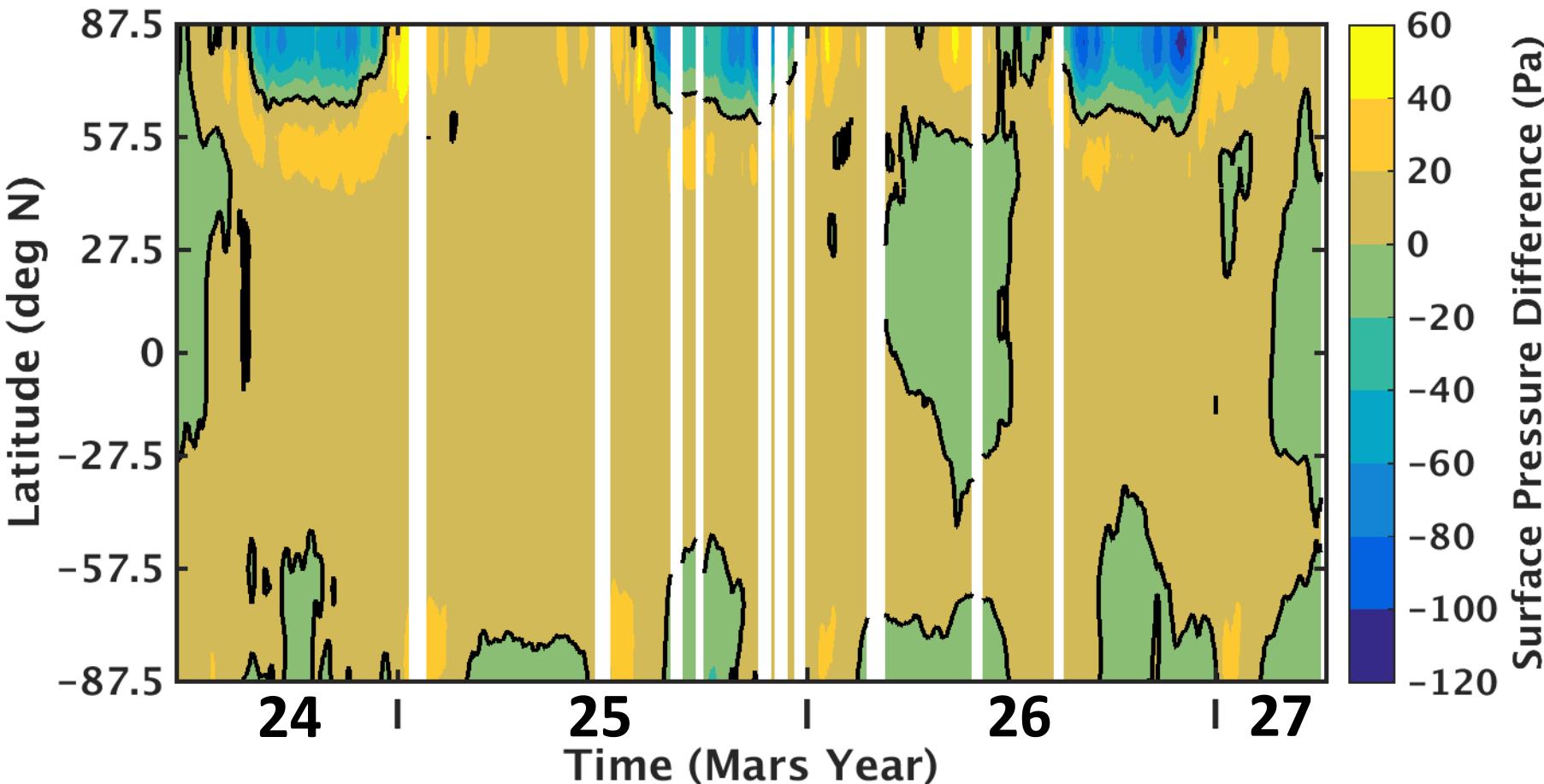
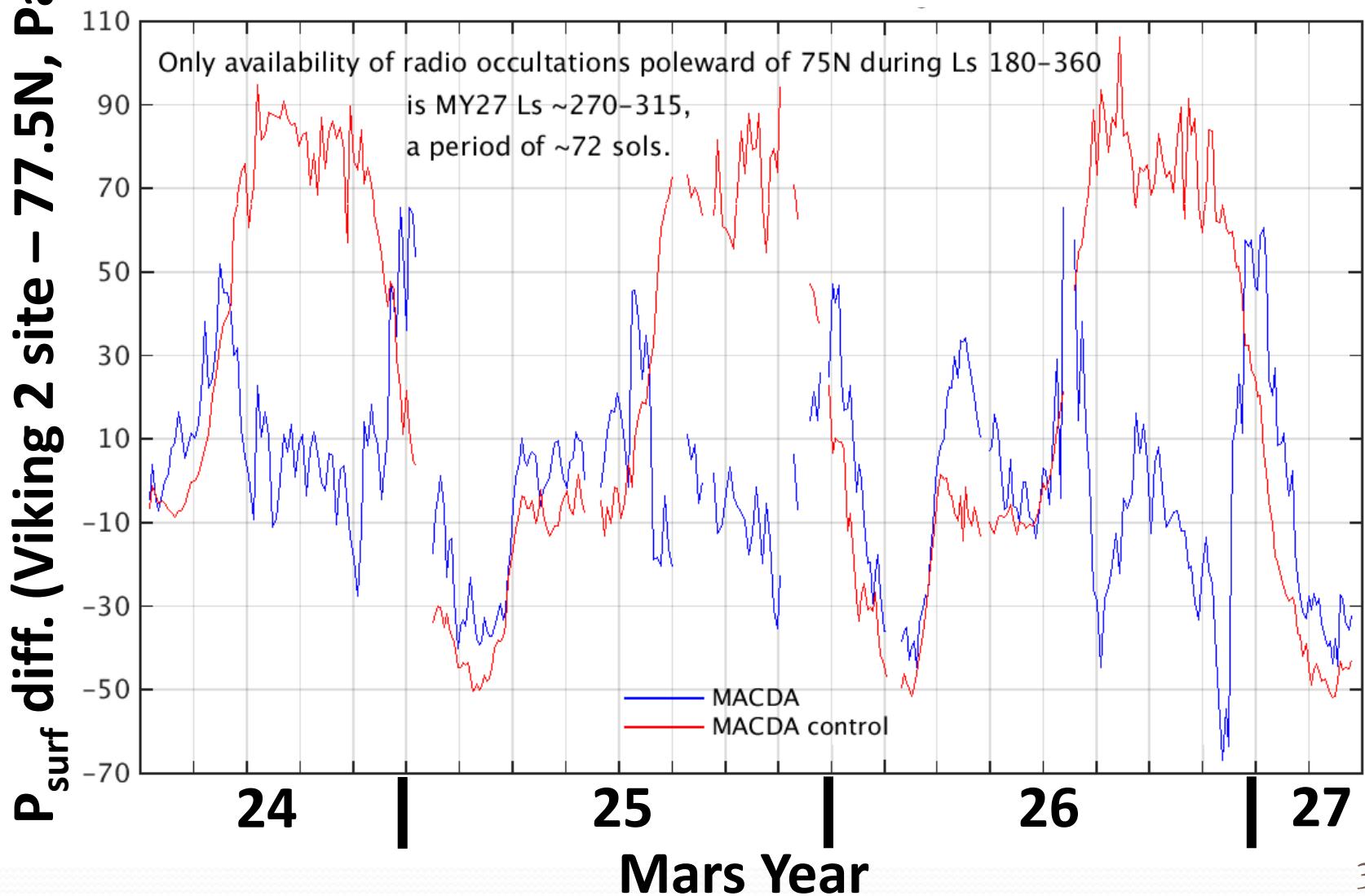


Figure: G. Davis

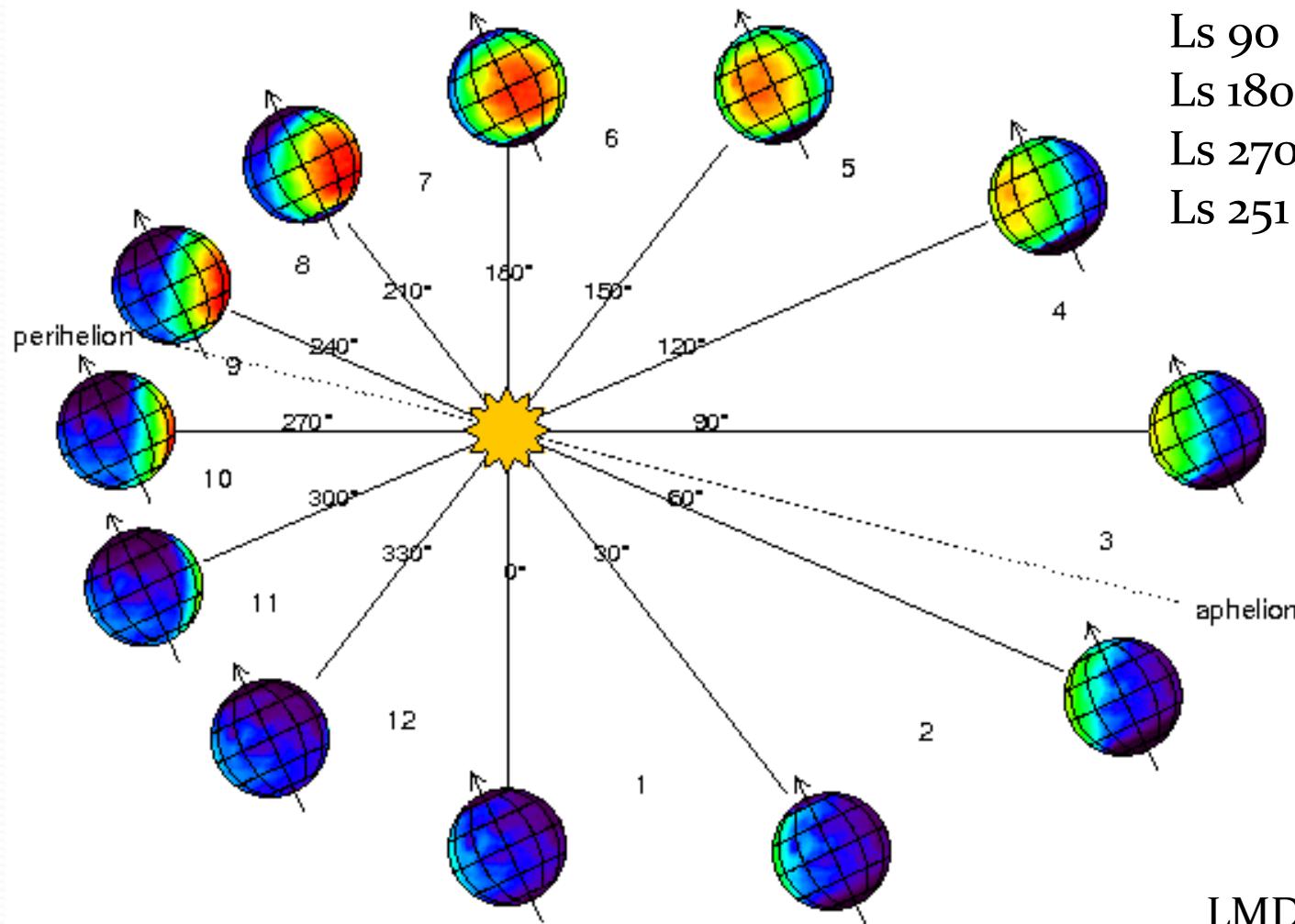
P_{surf} difference may be observable



(Mars) conclusions and caveats

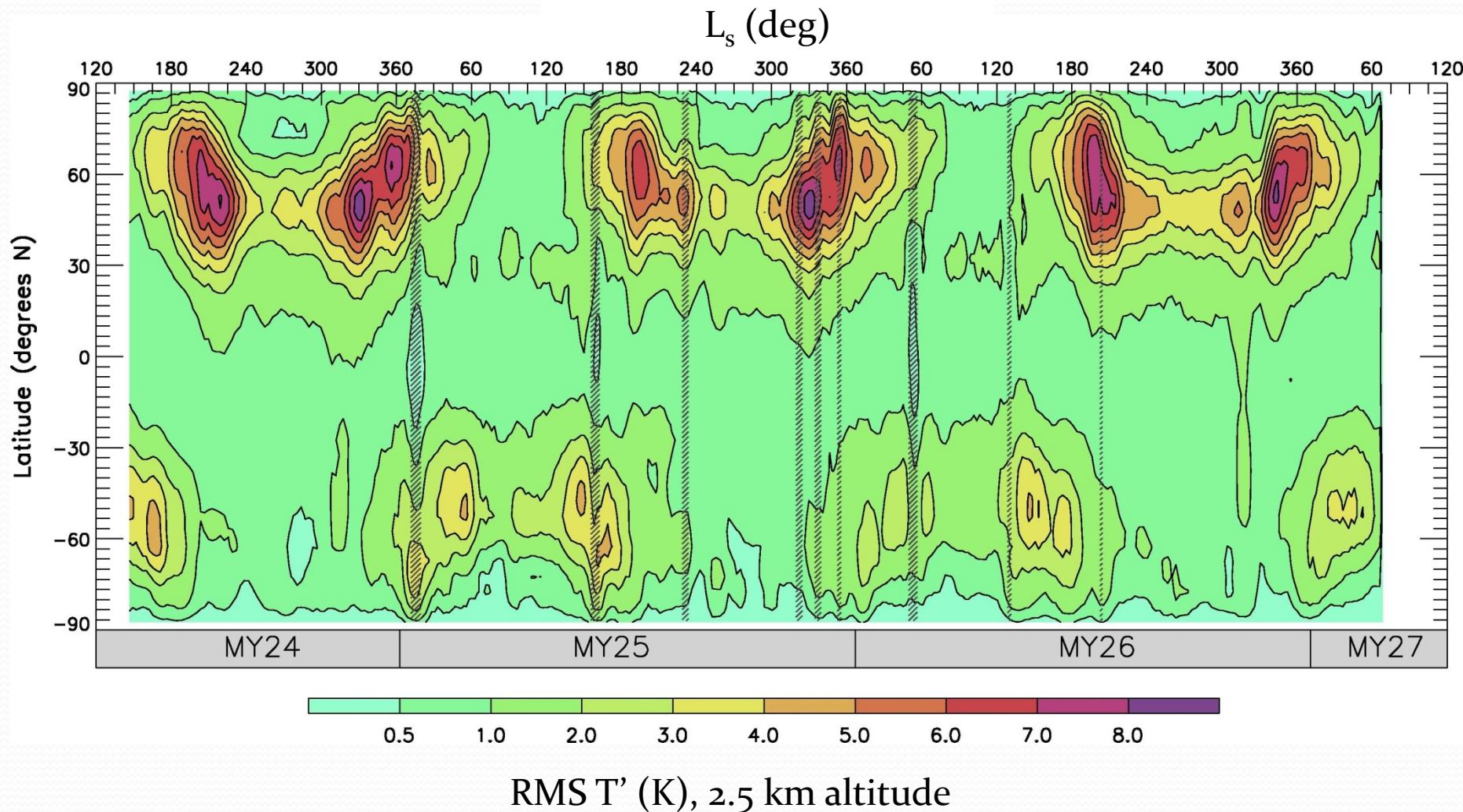
- Conclusions
 - Data assimilation notably alters low-level circulation
 - “Truth” not obvious, but could likely learn more about it without new Mars mission
- Caveats
 - Differences less stark in more modern reanalysis (EMARS)

Mars calendar



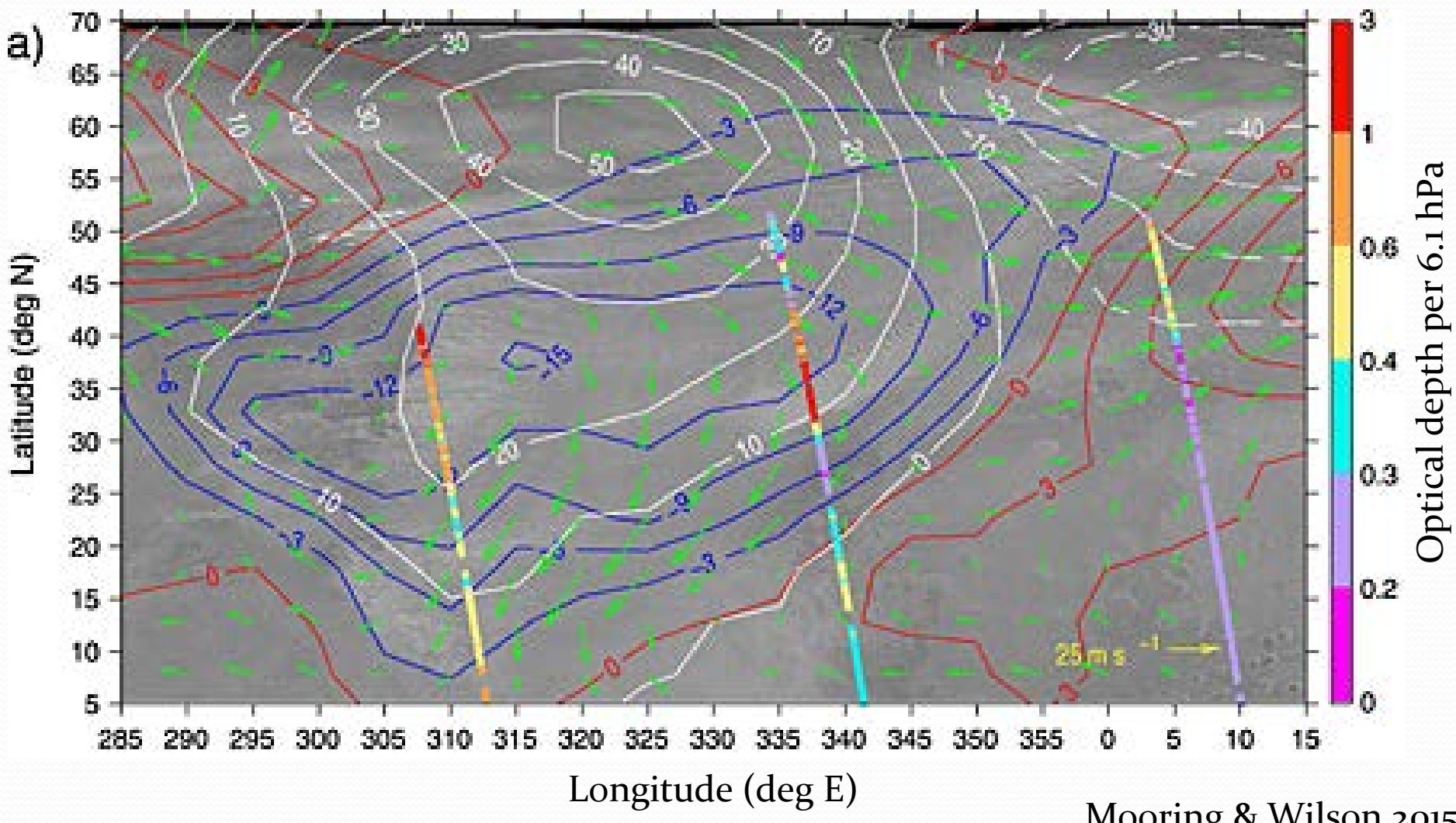
LMD Mars Climate
Database n.d.

Seasonal cycle of low-level eddies



Flushing dust storm in Acidalia/Chryse

Green: Total wind ($\tilde{A} = 0.9$, m/s) Red/blue: T' ($\tilde{A} = 0.9$, K) White: p_s' (Pa) MY24 L_s 224



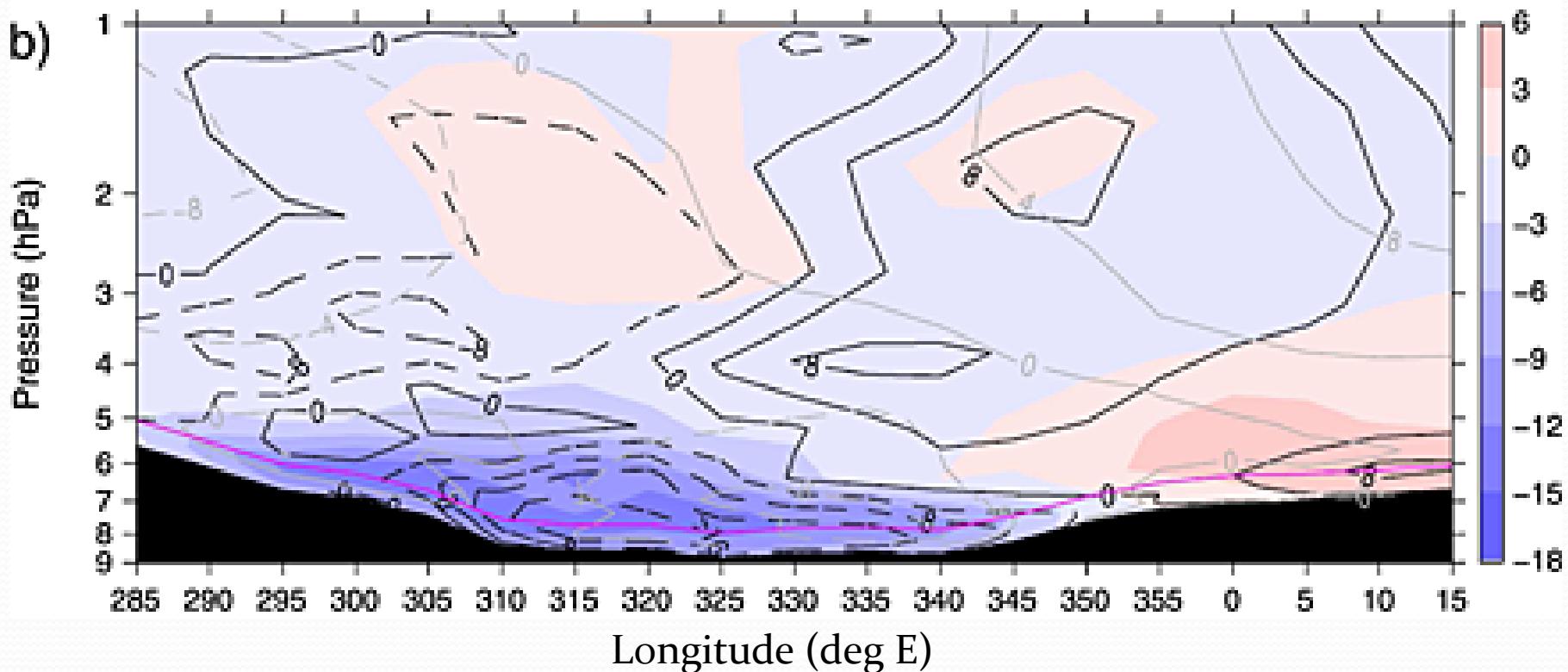
Flushing dust storm in Acidalia/Chryse

Black: v' (m/s)

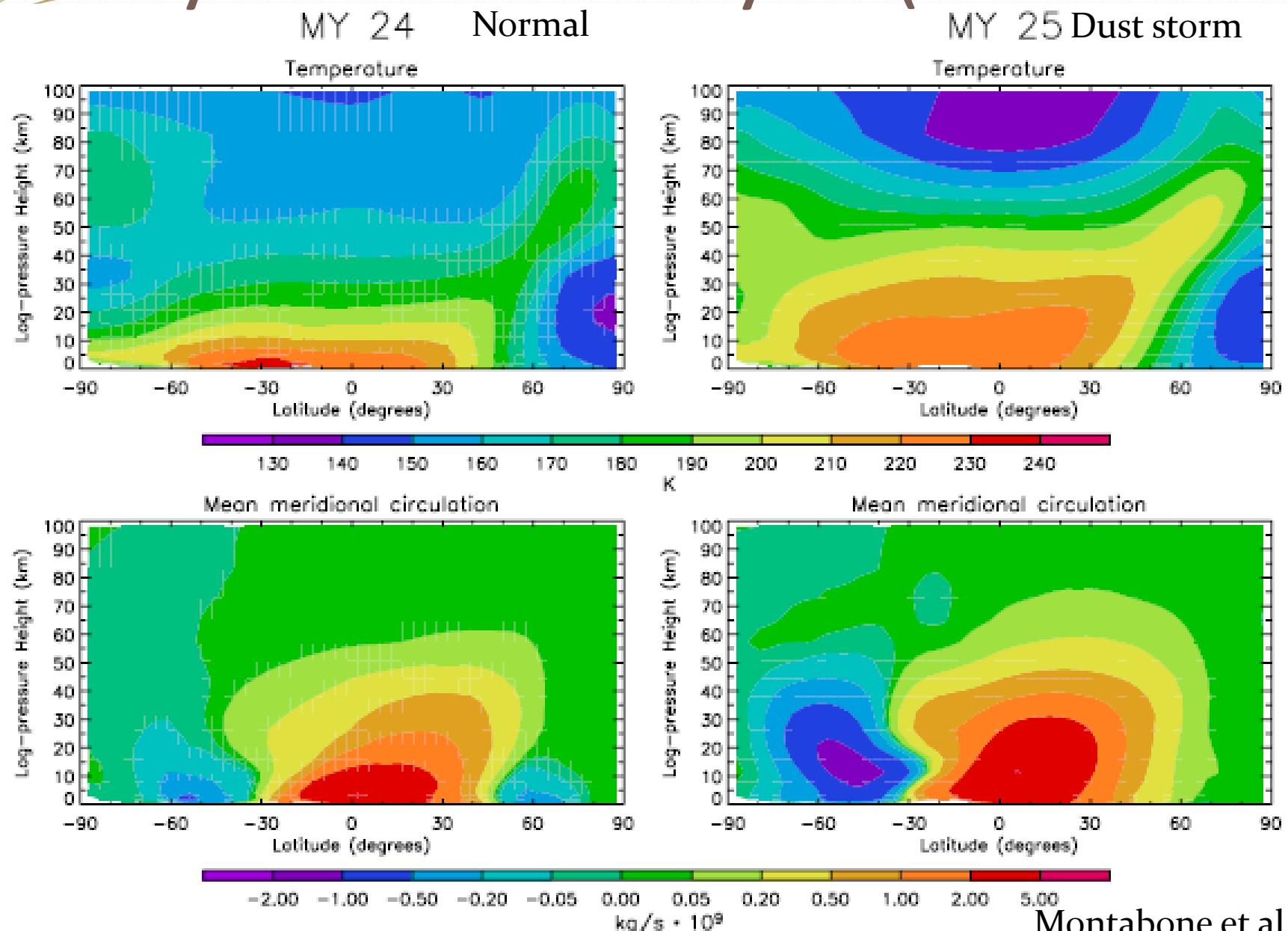
Red/blue: T' (K)

Gray: quasistationary v (m/s)

MY24 L_s 224, 32.5N



Hadley cell in reanalysis (Ls 195-225)



Mars topographic map

